## DESIGN FILE MEMORANDUM

DATE: February 13, 2001
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SUBJECT: ARP T002_0005 release
FILENAME: sdfm240_arpt002_0005.fm

The MSR OBC has been used to detect degradations of the cameras since launch. Before this could be done, however, the OBC had to itself be calibrated. This has been done, and documented in SDFM\#231, MISR On-Board Calibrator performance, N. Chrien. This memo documents the corrections constants, k , that we believe need to be applied to the OBC photodiodes in order to make them each consistent with the HQE_blue photodiode.

$$
\begin{equation*}
\mathrm{L}_{\text {measured }}^{\mathrm{std}, \mathrm{OBC}}=\frac{1.2395 \mathrm{iE}_{0}^{\mathrm{std}}}{(\mathrm{~A} \Omega \Re) \cdot \mathrm{k}} \tag{1}
\end{equation*}
$$

This factor of $k$ has been found to be variant in time, and has the values:
Table 1. OBC correction factors based upon the HQE_blue photodiode (SDFM\#231)

| Orbit | Notes | Band |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
|  |  | BLUE | GREEN | RED | NIR |
| 1076 |  | 1.0000 | 1.0405 | 0.9575 | 1.0955 |
| 1259,1912 |  | 1.0000 | 1.0338 | 0.9571 | 1.0793 |
| 1314,1912 |  | 1.0000 | 1.0337 | 0.9573 | 1.0782 |
| 1911,1912 | April 27. Used for <br> ARP 4 | 1.0000 | 1.0300 | 0.9561 | 1.0682 |
| 2575,2585 | June 11 | 1.0000 | 1.0257 | 0.9536 | 1.0531 |
| 3717,3721 |  | 1.0000 | 1.0212 | 0.9538 | 1.0330 |

Using these coefficients, the following radiometric response coefficients were reported for the AN cameras. (These data were read from the files available in /data/ifrcc/workspace/out/ production.)

Table 2. G1 coefficients

| Date | BLUE | GREEN | RED | NIR |
| :--- | :--- | :--- | :--- | :---: |
| April 27, 2000 | 22.5434 | 22.9652 | 30.7784 | 45.4112 |
| June 11, 2000 | 22.3108 | 22.5464 | 29.1454 | 42.7529 |

We now wish to update this table, based upon the MISR Vicarious Calibration experiment, dated June 11th. These results have shown that:

Table 3. Ratio of MISR/ VC radiances

| Parameter | BLUE | GREEN | RED | NIR |
| :--- | :--- | :--- | :--- | :--- |
| Measured (Ver 4) | 0.89 | 0.88 | 0.93 | 0.92 |
| Adjusted per Equation 3 | 0.899 | 0.896 | 0.982 | 0.977 |
| Used for ARP processing | 0.90 | 0.98 |  |  |

That is, we wish to increase the radiances reported by MISR on June 11th by the inverse of these numbers. We first note that that comparison made in Table 3 was done using the MISR gain coefficients reported in ARP Version 4 (based upon April 27th data). We would have liked to have had available ARP coefficients based upon the June 11th experiment. We can make this adjustment by noting

$$
\begin{equation*}
\mathrm{L}_{\text {misr,Jun11 }}=\mathrm{L}_{\text {misr,arp4 }}\left(\frac{\mathrm{G}_{1, \text { ARP4 }}}{\mathrm{G}_{1, \text { Jun11 }}}\right) \tag{2}
\end{equation*}
$$

therefore the actual ratio we are interested in is

$$
\begin{equation*}
\mathrm{k}_{\mathrm{vc}}{ }^{\prime}=\frac{\mathrm{L}_{\text {misr,Jun11 }}}{\mathrm{L}_{\mathrm{vc}}}=\mathrm{k}_{\mathrm{vc}}\left(\frac{\mathrm{G}_{1, \text { ARP4 }}}{\mathrm{G}_{1, \text { Jun } 11}}\right) \tag{3}
\end{equation*}
$$

To a new algorithm to compute the photodiode reported radiances:

$$
\begin{equation*}
\mathrm{L}_{\text {measured }}^{\mathrm{std}, \mathrm{OBC}}=\frac{1.2395 \mathrm{iE}_{0}^{\mathrm{std}}}{(\mathrm{~A} \Omega \Re) \cdot \mathrm{k}_{231}}\left(\frac{\mathrm{G}_{1, \mathrm{ARP} 4}}{\mathrm{G}_{1, \text { Jun } 11} \mathrm{k}_{\mathrm{vc}}}\right) \tag{4}
\end{equation*}
$$

where
$\mathrm{k}_{231} \quad$ are the constants given in Table 1
$\mathrm{G}_{1, \text { ARP4 }} \quad$ are the constants given in Table 2, row 1
$\mathrm{G}_{1, \text { Jun11 }} \quad$ are the constants given in Table 2, row 2
$\mathrm{k}_{\mathrm{vc}}$ are the constants given in Table 3, row $3=\{0.90,0.90,0.98,0.98\}$

We therefore see that we may still use the procedure developed in SDFM\#231, but with a new diode constant k:

$$
\begin{equation*}
\mathrm{k}_{\text {diode }}=\mathrm{k}_{213}\left(\frac{\mathrm{G}_{1, \text { Jun11 }} \mathrm{k}_{\mathrm{vc}}}{\mathrm{G}_{1, \text { ARP4 }}}\right)=\mathrm{k}_{213} \mathrm{k}_{\mathrm{vc}}{ }^{\prime} \tag{5}
\end{equation*}
$$

